

PHZ 6607 Special and General Relativity October 3, 2018

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1 Scattering in Schwarzschild

Today we will discuss unbound orbits in the Schwarzschild geometry, or scattering off of a Schwarzschild black hole. Previously, we derived the geodesic equation

$$\left(\frac{dr}{d\tau}\right)^2 = c^2((E/mc^2)^2 - 1) + \frac{2GM}{r} - \frac{L^2}{m^2r^2} + \frac{2GML^2}{m^2r^3c^2},$$

where $E = m(1 - 2GM/rc^2)dt/d\tau$ is the relativistic energy of the particle at infinity and $L = mr^2 \sin^2 \theta d\phi/d\lambda$ is the angular momentum of the particle. We also defined the variable $u = 1/r$, the angular momentum per unit mass $\tilde{L} = L/m$, and the relativistic energy per unit mass $\tilde{E} = E/mc^2$. Making these substitutions, we found

$$\frac{d^2u}{d\phi^2} = \frac{GM}{\tilde{L}^2} - u + \frac{3GM}{c^2}u^2$$

We defined the impact parameter b by $GM/\tilde{L}^2 = 1/b$. We will now solve this equation to first order in the quantity $3GM/c^2$, so we write

$$u = u_0 + \left(\frac{3GM}{c^2}\right)u_1 + O\left(\left(\frac{3GM}{c^2}\right)^2\right)$$

1.1 Zeroth order

The zeroth order equation is then

$$\frac{d^2u_0}{d\phi^2} = \frac{1}{b} - u_0.$$

The solution to this equation is

$$u_0 = \frac{1}{b}(1 + \varepsilon \cos \phi),$$

where $\varepsilon > 1$ is the eccentricity for a hyperbolic orbit. Recall that

$$\pi/2 < \phi_{\max} = \arccos(-1/\varepsilon) < \pi.$$

Taking the derivative, we get

$$\frac{du_0}{d\phi} = -\frac{\varepsilon}{b} \sin \phi = -\frac{1}{r^2} \frac{dr}{d\phi}.$$

Since $du/d\phi = -(1/r^2)dr/d\phi$, we can now examine what happens at large r , or as $\phi \rightarrow \phi_{\max}$. Taking this limit,

$$\lim_{\phi \rightarrow \phi_{\max}} \frac{1}{r^2} \frac{dr}{d\phi} = \frac{\epsilon}{b} \sin(\arccos(-1/\epsilon)) = \frac{\epsilon}{b} \sqrt{1 - (1/\epsilon)^2} = \frac{\sqrt{\epsilon^2 - 1}}{b} = \sqrt{\frac{c^2}{\tilde{L}^2} (\tilde{E}^2 - 1)},$$

Separately, since in this limit we know that

$$\frac{dr}{d\tau} = \frac{\tilde{L}}{r^2} \frac{dr}{d\phi} = c \sqrt{\tilde{E}^2 - 1},$$

we can solve for the eccentricity:

$$\begin{aligned} \frac{\sqrt{\epsilon^2 - 1}}{b} &= \sqrt{\frac{c^2}{\tilde{L}^2} (\tilde{E}^2 - 1)} \\ \implies \epsilon^2 &= 1 + \left(\frac{\tilde{L}c}{GM} \right)^2 (\tilde{E}^2 - 1). \end{aligned}$$

This reproduces the non-relativistic result.

2 First order

Given knowledge of u_0 , the first order equation is

$$\frac{d^2 u_1}{d\phi^2} = -u_1 + \frac{3GM}{c^2} \frac{1}{b^2} (1 + 2\epsilon \cos \phi + \frac{\epsilon^2}{2} (1 + \cos 2\phi)).$$

The solution to this equation is given by

$$u_1 = A + B\phi \cos \phi + c \cos 2\phi,$$

where

$$A = \frac{3GM}{c^2} \frac{1}{b^2} (1 + \epsilon^2/2), \quad B = \frac{3GM}{c^2} \frac{1}{b^2} \epsilon, \quad C = -\frac{GM}{c^2} \frac{1}{b^2} \frac{\epsilon^2}{2}.$$

Thus, our solution for $u = 1/r$ to first order is

$$u = u_0 + u_1 = \frac{1}{b} \left(1 + \frac{GM}{bc^2} (3 + 2\epsilon^2) + \epsilon \cos \left(\left(1 - \frac{3GM}{bc^2} \right) \phi \right) - \frac{GM}{bc^2} \epsilon \cos^2 \phi \right).$$

By combining the trigonometric terms and taking an approximation, we arrive at the new deflection angle

$$\phi_{\max} = \left(1 - \frac{3GM}{bc^2} \right)^{-1} \arccos \left(-\frac{1 + \frac{2GM}{bc^2} (1 + \epsilon^2)}{\epsilon} \right),$$

which is advanced compared to the ϕ_{\max} in the Newtonian case.